# AERONOMICA ACTA 

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A-N^{\circ} 327-1988
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European validation of Sage II aerosol profiles

> by
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## FOREWORD

The article "European Validation of SAGE II Aerosol Profiles" will appear in the Journal of Geophysical Research.

AVANT-PROPOS

L'article "European Validation of SAGE II Aerosol Profiles" sera publié dans le Journal of Geophysical Research.

VOORWOORD

Het artikel "European Validation of SAGE II Aerosol Profiles" zal in het Journal of Geophysical Research gepubliceerd worden.

VORWORT

Der Artikel "European Validation of SAGE II Aerosol Profiles" wird im Journal of Geophysical Research publiziert werden.

European validation of sage il aerosol profiles

## by

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## Abstract

A SAGE II validation programme has been performed in Europe using ground-based lidars and balloon borne polarimetric and photographic experiments. Between the tropopause height and about 23 km , a good agreement is found between the SAGE II $1.02 \mu \mathrm{~m}$ extinction profiles and the lidar profiles, using for the conversion of backscattering into extinction and aerosol model consistent with the SAGE II spectral extinction. The extinction profiles deduced from the limb photographs at $.44 \mu \mathrm{~m}$ and $.375 \mu \mathrm{~m}$ present a good agreement with the SAGE II profiles respectively at $.453 \mu \mathrm{~m}$ and $.385 \mu \mathrm{~m}$. The size distribution retrieved from the near infrared polarimetric observations leads to a spectral variation of the extinction in good agreement with SAGE II data, in the same altitude range. Above $23-25 \mathrm{~km}$, the observations are scarce and the data of poorer quality because of the low aerosol content. The $1.02 \mu \mathrm{~m}$ extinction profiles seem to agree with the ruby lidar and the limb photographs profiles. But any conclusion concerning the short wavelength profiles and the size distribution at this high altitudes would be risky.

Un programme de validation de l'instrument en orbite "SAGE II" a été exécuté en Europe utilisant des lidars au sol et des expériences de polarimétrie et photographiques par ballons stratosphériques. Entre l'altitude de la tropopause et environ 21 km un bon accord est obtenu entre les profils d'extinction à $1.02 \mu \mathrm{~m}$ de SAGE II et les profils lidar, utilisant pour la conversion de la rétrodiffusion en extinction un modèle d'aérosol en accord avec l'extinction spectrale de SAGE II. Les profils d'extinction déduits des photographies du limbe terrestre à $.44 \mu \mathrm{~m}$ et $.375 \mu \mathrm{~m}$ présentent un bon accord avec les profils de SAGE II à $.453 \mu \mathrm{~m}$ et $.385 \mu \mathrm{~m}$ respectivement. La distribution de taille des particules déduites des observations polarimétrique infrarouge proche conduisent à une dépendance spectrale de l'extinction en bon accord avec SAGE II dans le même domaine d'altitudes. Au-dessus de 23-25 km les observations sont moins nombreuses et les*données de moins bonne qualite par suite du faible contenu en aerosol. Les profils d'extinction à $1.02 \mu \mathrm{~m}$ semble être en accord avec les profils obtenus par lidar à rubis et par photographies. Cependant toute conclusion concernant les profils aux courtes. longueurs d'onde et la distribution de taille à ces altitudes élevées seraient risquées.

Een validatieprogramma van het instrument "SAGE II" werd in Europa uitgevoerd, waarbij gebruik gemaakt werd van grondlidars en ballongedragen polarimetrische en fotografische experimenten. Tussen de hoogte van de tropopauze en ongeveer 23 km is een goede overeenstemming gevonden tussen de SAGE II $1.02 \mu \mathrm{~m}$ uitdovingsprofielen en de Iidarprofielen, waarbij voor de omzetting van de retrodiffusie in uitdoving een aërosolmodel gebruikt werd dat overeenstemt met de spectrale uitdoving van SAGE II. De uitdovingsprofielen afgeleid uit de foto's van de aardse kim op $.44 \mu \mathrm{~m}$ en $.375 \mu \mathrm{~m}$ stemmen goed overeen met de. SAGE II profielen op respectievelijk . $453 \mu \mathrm{~m}$ en $.385 \mu \mathrm{~m}$. De verdeling volgens grootte van de deeltjes afgeleid uit de polarimetrische waarnemingen in het nabije infrarood leidt tot een spectrale afhankelijkheid van de uitdoving die goed overeenstemt met SAGE II in hetzelfde noogtegebied. Boven $23-25 \mathrm{~km}$ zijn de waarnemingen schaars en de gegevens van minder goede kwaliteit als gevolg van de geringe aërosolaanwezigheid. De uitdovingsprofielen op $1.02 \mu \mathrm{~m}$. schijnen overeen te stemmen met de profielen bekomen met behulp van de robijnlidar en foto's. Elk besluit omtrent de profielen voor korte golflengten'en de verdeling volgens grootte zou bij deze grote hoogten echter gewaagd zijn.

Ein Validationprogramm des Instrumentes "SAGE II" wurde in Europa durchgeführt, wobei Grundlidars und ballongetragen polarimetrische und photografische Experimenten gebraucht wurden. Zwischen die Höhe der Tropopause und ungefähr 23 km ist eine gute Übereinstimmung gefunden zwischen den. SAGE II $1.02 \mu \mathrm{~m}$ Extinktionprofilen und den Lidarprofilen, wobel für die Umsetzung der Retrostreuung in Extinktion ein Aerosolmodell gebraucht wurde in Ubereinstimmung mit der spektralen Extinktion von SAGE II. Die Extinktionprofilen abgeleitet von Photografien des irdischen Horizontes am . 44 . $\mu \mathrm{m}$ und. $375 \mu \mathrm{~m}$ stimmen gut überein mit den SAGE II Profilen am respektive . $453 \mu \mathrm{~m}$ und $.385 \mu \mathrm{~m}$. Die Grösseverteilung der Partikeln abgeleitet von polarimetrischen Beobachtungen im nahe Infrarot leitet zu einer spektralen Abhängigkeit der Extinktion die gut übereinstimmt mit SAGE II im gleichen Höhegebiet. Über 23-25 km sind die Beobachtungen selten und die Daten von schlechterer Qualität infolge die geringe Aerosolanwesenheit. Die Extinktionprofilen am $1.02 \mu \mathrm{~m}$ scheinen uberein zu stimmen mit den Profilen bekommen mit der. Hilfe vom Rubinlidar und Photografien. Aber jeder Schluss über den Profilen für kurze Wellenlängen und der Grösseverteilung lst auf dieser grossen Höhen gewagt.

SAGE II provides aerosol extinction profiles at $1.02 \mu \mathrm{~m}$, $.525 \mu \mathrm{~m}$, $.453 \mu \mathrm{~m}$ and $.385 \mu \mathrm{~m}$. The $1.02 \mu \mathrm{~m}$ channel is free from any other contribution except the Rayleigh correction, which is not too large for most of the altitude range; the results are retrieved without difficulties almost down to the ground level in cloud free cases. The upper limit of retrieval is due to the low aerosol concentration which leads to transmissions very close to one above some altitude, which is around 30 km at middle latitudes, for the present state of the atmosphere; smoothing procedures allow the retrieval at higher altitudes but with increasing error bars. The three short wavelength channels are contaminated by ozone and nitrogen dioxide absorption; the separation of these contributions has been discussed in Chu et al, (same issue). The upper limit of a reasonable quality retrieved profile is probably a little below 30 km for the three channels, because of the increasing contribution of $\mathrm{O}_{3}$ and $\mathrm{NO}_{2}$ with altitude, Moreover the Rayleigh correction increases toward the short wavelengths leading to an increasing error on the retrieved extinctions and limiting the retrieval to altitudes above 8,10 and 14 km for the $.525 \mu \mathrm{~m}, .453 \mu \mathrm{~m}$ and $.385 \mu \mathrm{~m}$ aerosol extinction profiles, respectively.

The aerosol extinction depends on their total number density and on their size distribution, as well as their shape and refractive index; they are generally assumed to be spherical droplets of an aqueous sulfuric acid solution. This means that there is no direct, simple validation experiment for the aerosol data, as there is for the gas data, where only one parameter (the gas concentration) has to be measured. The most direct approach is to measure in situ the absolute size distribution $n(r)$ of the particles (including the total number $N=$ $\int_{0}^{\infty} n(r) d r$ per unit volume) and to compute by Mie theory extinction profiles to be compared to the SAGE II extinction profiles. This can be achleved by various instruments, such as wire impactors, QCM, multifilters, optical counters (Russel.l et al., 1981, Russel.1 et al., 1984,

Oberbeck et al., 1986); however all instruments have limitations in the range of sizes detected. Another approach consists in using other scattering measurements which should be consistent with the SAGE II extinction profiles. The best known example of this procedure is the lidar backscattering profile; the data have to be converted into extinction profiles, using a model of the aerosol size distribution and the Mie theory; for consistency the same model must reproduce the spectral variation of the extinction deduced from the four SAGE II aerosol channels.

During the correlative experiments made in Europe (Lenoble, same issue), ground based lidars were used at four different stations; the results are presented in section 2. Two other scattering techniques have been simultaneously used from balloon platforms : photographs of the earth's limb and infrared polarimetric measurements; these are presented with some details in sections 3 and 4 respectively. Analysis and comparisons of the results are discussed in section 5 .

## 2. LIDAR

Nd Yag lasers operating at . 532 um have been used at Observatoire de Haute Provence (OHP), Frascatti and Florence, and a ruby laser operating at . $694 \mu \mathrm{~m}$ at Garmisch-Partenkirchen; the ruby laser allows retrieval of the profiles up to altitudes higher than the Nd Yag laser, because of the smaller contribution of Rayleigh scattering at the larger wavelength. The data provided by the experimenters are the backscattering ratio $R=\left(b^{\text {aer }}+b^{m o l}\right) / b^{\text {mol }}$, versus the altitude; $b^{\text {aer }}$ is the aerosol and $b^{m o l}$ the molecule backscattering coefficient.

The aerosol backscattering is deduced from $R$ and $b^{\text {mol }}$ (computed for a standard atmosphere). Then the aerosol extinction coefficient at the laser wavelength derives by

$$
\begin{equation*}
\sigma^{\text {aer }}(\lambda)=4 \pi \text { b }^{\text {aer }} / \mathrm{p}^{\text {aer }}\left(\lambda, 180^{\circ}\right) \tag{1}
\end{equation*}
$$

where the phase function $p^{\text {aer }}$ is computed from an aerosol model; the same model is used to transform $\sigma^{\text {aer }}(\lambda)$ into $\sigma^{\text {aer }}(1.02)$ for comparison with the SAGE II profiles at $1.02 \mu \mathrm{~m}$. We have chosen to compare the IIdar data with the SAGE II profiles at $1.02 \mu \mathrm{~m}$, because they are of better quality and retrieved at higher altitudes than the SAGE II profiles at $.525 \mu \mathrm{~m}$, which are closer to the lidar wavelengths. Similarly the statistical error $\Delta R$ due to signal fluctuations is transformed into an error $\Delta 0^{\text {aer }}(1.02)$. The conversion of backscattering at $\lambda$ into extinction at $1.02 \mu \mathrm{~m}$ has been made with various models and it has proved insensitive to the detailed shape of the size distribution, the main parameter being the effective radius $r_{\text {eff }}$ and the effective variance $v_{\text {eff }}$ defined by

$$
\begin{align*}
& r_{e f f}=\int_{0}^{\infty} r^{3} n(r) d r / \int_{0}^{\infty} r^{2} n(r) d r  \tag{2}\\
& v_{e f f}=\int_{0}^{\infty}(r-r e f f)^{2} r^{2} n(r) d r / r_{e f f}^{2} \int_{0}^{\infty} r^{2} n(r) d r \tag{3}
\end{align*}
$$

The simplest choice for modeling is a log-normal (LND) size distribution

$$
\begin{equation*}
n(r)=\frac{N}{\sqrt{2 \pi} r \ln \sigma} \exp \left[-\frac{\ln ^{2} r / r_{m}}{2 \ln ^{2} \sigma}\right] \tag{4}
\end{equation*}
$$

$r_{\text {eff }}$ and $v_{\text {eff }}$ are related to the mean radius $r_{m}$ and the variance $\sigma$ by (Lenoble and Brogniez, 1984)

$$
\begin{align*}
& r_{\text {eff }}=r_{m} \exp \left(2.5 \ln ^{2} \sigma\right)  \tag{5}\\
& v_{\text {eff }}=\exp \left(\ln ^{2} \sigma\right)-1 \tag{6}
\end{align*}
$$

Figure 1 presents the conversion factor $K(\lambda)=\sigma^{\text {aer }}(\lambda) p^{\text {aer }}\left(\lambda, 180^{\circ}\right) /$ $\sigma^{\text {aer }}(1.02)$ versus $r_{\text {eff }}$ for LND models $\left(75 \% \mathrm{H}_{2} \mathrm{SO}_{4}\right.$ droplets at 220 K$)$ with $v_{\text {eff }}=.25(\sigma=1.60)$, for the two lidars. It varies slowly for $r_{\text {eff }}$ larger than $.20 \mu \mathrm{~m}$ and is almost constant for $\mathrm{r}_{\text {eff }}$ larger than $.30 \mu \mathrm{~m}$; but it increases very rapidly for smaller particles. Similar curves can be drawn for other values of $v_{\text {eff }}$; however the influence of $v_{\text {eff }}$ becomes larger only for particles smaller than $.20 \mu \mathrm{~m}$; it is almost negligible for $r_{\text {eff }}=.30 . \mu \mathrm{m}$, as K increases with $\mathrm{v}_{\mathrm{eff}}$ for $\mathrm{r}_{\mathrm{eff}}>.30 \mu \mathrm{~m}$ and decreases for $r_{\text {eff }}<.30 \mu \mathrm{~m}$.

The results of the lidar/SAGE II profile comparisons will be presented, using for conversion of backscattering into extinction the model (or models) with $v_{\text {eff }}=.25$, which gives the best agreement between the two profiles. The choice will be checked to be consistent with the SAGE II spectral extinction, and/or with the in situ observations, in section 5.

The four periods of observations were November 10-13, 1984, November 27-30, 1984, April 21-23, 1985 and October 12-14, 1985; the location of the SAGE II events and of the ground stations are presented in Lenoble (same issue).

During the period November 10-13, 1984, the aerosol layer was very unstable on the local scale, as shown by the in sitid balloon observations, and on the scale of the observation zone, as proved by the important differences between the six SAGE II profiles. The lidar profiles obtained at the OHP and at Garmisch-Partenkirchen on November 11 are also quite different, as are the profiles obtained at the OHP on November. 11 and November 13. However on November 13, the situation seems to stabilize and the two SAGE II profiles at 8049 W ( 1703 GMT) and $15^{\circ} 81 \mathrm{E}$ ( 1527 GMT) are very similar; this allows a comparison with the OHP lidar profile obtained at the same latitude and $6^{\circ} \mathrm{E}$, from 1705 to 1817 GMT (Figure 2). The agreement is very good above 18 km and certainly within the error limits of the lidar profile; below


Fig. 1.- Lidar conversion factor $K(\lambda)=\sigma^{\mathrm{aer}}(\lambda) \mathrm{p}^{\mathrm{aer}}\left(\lambda, 180^{\circ}\right) / \sigma^{\mathrm{aer}}(1.02)$ versus the effective radius $\mathrm{r}_{\text {eff }}$ for aerosol LND models with $v_{\text {eff }}=.25$; particles are $\mathrm{H}_{2} \mathrm{SO}_{4} 75 \%$ at 220 K ; solid curve for ruby lidar; dashed curve for Nd Yag lidar.


Fig. 2.- Comparison of aerosol extinction profiles at $1.02 \mu \mathrm{~m}$ from SAGE II and retrieved from lidar backscattering with a LND model $\mathrm{v}_{\mathrm{eff}}=.25$ (see text) on November 13, 1984.

- sojid curve : SAGE II, 1703 GMT, $44062 \mathrm{~N}, 8049 \mathrm{~W}$,
- dásh-dotted curve : SAGE II, 1527 GT, $44^{\circ} 50 \mathrm{~N}, 15^{\circ} 81 \mathrm{E}$ (error bárs ommited),
. dashed curve : lidar OHP, 1705 to 1817 GMT, $44^{\circ} \mathrm{N}, 6^{\circ} \mathrm{E}$ (aerosol model : $\mathrm{r}_{\text {eff }}=.35 \mu \mathrm{~m}$ ).

18 km the differences are probably due to the variability of the aerosol layer. The conversion from backscattering into extinction has been made with a LND model, $r_{\text {eff }}=.35 \mu \mathrm{~m}$, for all altitudes. But as mentionned above the conversion is not very sensitive to the effective radius in this size range and any model with $r_{\text {eff }}$ between. $20 \mu \mathrm{~m}$ and $.50 \mu \mathrm{~m}$ would lead to a similar agreement.

For the period November $26-30,1984$, the situation was more stable. From the eight SAGE II profiles, only the western profiles of November 28 and 29 seem to correspond to a different air mass with more aerosols above 25 km ; the other six profiles (eastern profiles for the whole period and western profiles for the 26 and 27) are very similar. The several jidar profiles obtained during this period (UHP, November 27-28-29 - Frascatti, November 28-29-30 - Garmisch-Partenkirchen, November 27) confirm a good stability of the aerosol layer over the zone and the period. This provides conditions much better than during the middle of November for a comparison programme. Figures 3, 4 and 5 present the results for these comparisons. For November 27 (Fig. 3) two lidar profiles are available, one from $O H P$ up to 25 km and one from Garmisch-Partenkirchen up to 30 km . To obtain agreement with SAGE II, the conversion from backscattering into extinction has to be made using models with a particle size decreasing with altitude $z$; we have used $r_{\text {eff }}=.25 \mu \mathrm{~m}$ for $z<21 \mathrm{~km}<\mathrm{z}<25 \mathrm{~km}$ and $r_{\text {eff }}=.05 \mu \mathrm{~m}$ for $\mathrm{z}>25 \mathrm{~km}$; a gradual change of $r_{\text {eff }}$ with altitude is more likely than an abrupt one, but should appear around $20-22 \mathrm{~km}$ and $25-26 \mathrm{~km}$. Above 28 km , even smaller particles would give better agreement than the model with $r_{\text {eff }}=$ $.05 \mu \mathrm{~m}$, but the accuracy of both the SAGE II and the lidar profiles is probably not good enough at these altitudes to give definite conclusion. The error bars on the Garmisch-Partenkirchen lidar profile have been given every two kilometers; they are somewhat larger for the OHP lidar profile. The two lidar profiles agree with each other reasonably well as well as with the SAGE II profile above 15 km ; below this they exhibit oscillations probably due to local conditions. For November 28 (Fig. 4) the conversion has been made with $r_{\text {eff }}=.25 \mu \mathrm{~m}$, up to 23 km , which is


Fig. 3.- Same as figure 2 on November 27, 1984,
. solid curve : SAGE II, 1457 GMT, $46^{\circ} 53 \mathrm{~N}, 18^{\circ} 68 \mathrm{E}$,
. dash-dotted curve : lidar OHP, $1720 \mathrm{GMT}, 44^{\circ} \mathrm{N}, 6^{\circ} \mathrm{E}$,
. dahsed-curve : lidar Garmisch-Partenkirchen, $47^{\circ} 5 \mathrm{~N}, 11^{\circ} \mathrm{E}$, (aerosol model : $r_{\text {eff }}=.25 \mu \mathrm{~m}, \mathrm{z}<21 \mathrm{~km} ; \mathrm{r}_{\text {eff }}=.10 \mu \mathrm{~m}, 21 \mathrm{~km}$ $\left.<z<25 \mathrm{~km} ; \mathrm{r}_{\mathrm{eff}}=.05 \mu \mathrm{~m}, \mathrm{z}>25 \mathrm{~km}\right)$.


Fig. 4. - Same as figure 2 on November 28, 1984, . solide curve : SAGE II 1509 GMT, $45^{\circ} 18 \mathrm{~N}, 16^{\circ} 76 \mathrm{E}$, . dash-double dotted curve : SAGE II, $1645 \mathrm{GMT}, 45^{\circ} 07 \mathrm{~N}, 7^{\circ} 36 \mathrm{~W}$ (error bars omitted),

- dashed curve : 1 idar OHP, 1657 to $1830 \mathrm{GMT}, 44^{\circ} \mathrm{N}, 6^{\circ} \mathrm{E}$, . dash-dotted curve : lidar Frascatti $1800 \mathrm{GMT}, 42^{\circ} \mathrm{N}, 13^{\circ} \mathrm{E}$, (aerosol model : $r_{\text {eff }}=.25 \mu \mathrm{~m}$ ).


Fig. 5.- Same as figure 2 on November 30, 1984,
. solitd curve : SAGE II $1533 \mathrm{GMT}, 41^{\circ} 40 \mathrm{~N}, 13^{\circ} 61 \mathrm{E}$,

- dashed curve : ljdar Frascatti $1936 \mathrm{GMT}, 42^{\circ} \mathrm{N}, 13^{\circ} \mathrm{E}$ (aerosol model : $\left.r_{\text {eff }}=.35 \mu \mathrm{~m}\right)$.
the upper limit of the lidar profiles (OHP and Frascati). The agreement between the OHP lidar and the SAGE II profiles is very good; the Frascati lidar leads to somewhat smaller extinctions below 20 km . November 30 (Fig. 5) corresponds to the closest coincidence, as the SAGE II tangent point is about 90 km from Frascatti. The conversion has been made with $r_{\text {eff }}=.35 \mu \mathrm{~m}$, and the agreement between the Frascatti lidar and the SAGE II profiles is almost perfect in the altitude range of the lidar profile ( $12-22 \mathrm{~km}$ ).

In April 1985, the five SAGE II aerosol profiles available over the zone are almost identical, pointing to very stable conditions for the aerosol layer. Unfortunately, the weather conditions did not permit Iidar observations, except at Garmisch-Partenkirchen on April 21. As the SAGE II tangent points were not very close to Garmisch-Partenkirchen on April 21, and considering the homogeneity mentioned above, we have chosen to make the comparison with the average SAGE II profile for the considered zone and period. The result of this comparison is presented on Figure 6; the bars on the SAGE II average profile are the standard deviations; they are of the same order as the errors bars on individual profiles; no error bars have been given for the lidar profile, but the errors can be expected of the same order as on Figure 3. The conversion of backscattering into extinction had to be made with $\mathrm{r}_{\text {eff }} .=.35 \mu \mathrm{~m}$ below 21 km and $r_{\text {eff }}=.10 \mu \mathrm{~m}$ above, in order to find the best agreement between the two profiles; the change of particle size seems rather abrupt between $20-22 \mathrm{~km}$. The agreement is within the error bars.

In October 1985 the six SAGE II profiles are rather similar, but not as similar as in April. Lidar profiles were obtained at OHP for four successive days and they show a good stability of the aerosol layer at this station. Figure 7 compares the OHP lidar profile on October 12 evening, with the SAGE II morning profile on October 12 approximately $10^{\circ} \mathrm{E}$ of OHP. The SAGE II profile at about $14^{\circ} \mathrm{W}$ of OHP is very similar above 16 km and so are the two profiles of October 13 morning, however with values slightly higher above 21 km for the eastern profile. Below


Fig. 6.- Same as figure 2 on April 21, 1985,
. solid curve : SAGE II, average of five profiles (April 21/23 between $50^{\circ} 14-43^{\circ} 80 \mathrm{~N}$ and $7^{\circ} 20 \mathrm{~W}-21^{\circ} 29 \mathrm{E}$ ) with standard deviations,
. dashed curve : lidar Garmisch-Partenkirchen, April $21,47^{\circ} 5 \mathrm{~N}$, $11^{\circ} \mathrm{E}$ (aerosol model : $r_{\text {eff }}=.35 \mu \mathrm{~m}, \mathrm{z}<21 \mathrm{~km} ; \mathrm{r}_{\mathrm{eff}}=.10 \mu \mathrm{~m}$, z > 21 km ).


Fig. 7.- Same as figure 2 on October 12, 1985,

- solid curve : SAGE II $0502 \mathrm{GMT}, 41^{\circ} 03 \mathrm{~N}, 17^{\circ} 03 \mathrm{E}$,
- dâshed curve : lidar OHP 1900 to $1920 \mathrm{GMT}, 44^{\circ} \mathrm{N}, 6^{\circ} \mathrm{E}$ (aerosol model $\left.r_{\text {eff }}=.17 \mu \mathrm{~m}\right)$,
- dash-dotted curve : same as dotted curve, with the aerosol model deduced from polarization measurements.

16 km the four profiles are quite different and no comparison can be sought. A good agreement above 16 km is found by using for the conversion of the lidar profile an aerosol model with $r_{\text {eff }}=.17 \mu \mathrm{~m}$.

In conclusion the extinction profiles retrieved from lidar profiles seem in reasonably good agreement (generally within the error bars) with the SAGE II extinction profiles at $1.02 \mu \mathrm{~m}$, provided $\mathrm{r}_{\text {eff }}$ is chosen suitably. The best validation is obtained on November 30, 1984, where there is a close coincidence in time and location between the SAGE II and Frascatti lidar profiles (Fig. 5). The periods of stability of the aerosol layer allow rather good validations with non coincident observations; this is the case of April 21-24, 1985 (Fig. 5). The conversion of backscattering lidar profiles into extinction profiles has to be done with variable models, the particle sizes being generally smaller at high altitudes. The consistency of the choice of the model with the other observations will be discussed in section 5 .

## 3. BALLOON LIMB PHOTOGRAPHS

Photographs of the limb radiance were made from balloon at Aire sur l'Adour by the IASB (Institut d'Aéronomie Spatiale de Belgique); the photographs were made for low sun elevation, at various solar azimuths for two wavelengths $.84 \mu \mathrm{~m}, .44 \mu \mathrm{~m}$ during the first flight, and for three wavelengths $84 \mu \mathrm{~m}, 44 \mu \mathrm{~m}$, and $.375 \mu \mathrm{~m}$ during the second flight (Ackerman et al., 1981). The extinction is deduced from the radiance measured at $30^{\circ}$ scattering angle, and the Rayleigh extinction is subtracted to obtain the aerosol extinction. Two flights took place on November 10, 1984 and April 22, 1985. They were simultaneous to flights of the polarimetric instrument described in section 4 .

On November 10, 1984, the photographs confirm the aerosol layer inhomogeneity mentioned previously (section 2). To the South of the balloon position, well marked vertical structures were observed (Ackerman et al, 1985). Towards the North a much smoother vertical
profile was observed. The SAGE II tangent point on November 10 is rather far to the South-East, over the Mediterranean Sea, and the SAGE II profiles on November 11 , closer to the bailoon launch site, are completely different than the November 10 profiles. We have therefore choosen to compare the balloon profiles with the two SAGE II profiles of November 10 and 11 in Figures 8 and 9. Figure 8 compares the balloon profile at $.44 \mu \mathrm{~m}$ with the two SAGE II profiles at. $453 \mu \mathrm{~m}$; we have not introduced a correction for the small wavelength difference. Figure 9 concerns the comparison at $1.02 \mu \mathrm{~m}$; the balloon profiles at $.84 \mu \mathrm{~m}$ have been converted to $1.02 \mu \mathrm{~m}$, using two aerosol models respectively with $r_{\text {eff }}=.28 \mu \mathrm{~m}$ and $r_{\text {eff }}=.05 \mu \mathrm{~m}$, but the conversion is not very sensitive to the model. At $1.02 \mu \mathrm{~m}$ the balloon profile on November 10 and the SAGE II profile on November 11 are in very close agreement, whereas the SAGE II profile on November 10 is quite different. At $.44 / .453 \mu \mathrm{~m}$, the balloon profile is between the two SAGE II profiles, with a general shape more similar to the SAGE II profile of November 11.

On April 22, 1985, the conditions were very stable, and the air mass observed at $30^{\circ}$ from the sun and for a tangent height of 20 km was very close to the SAGE II tangent point (Ackerman et al., 1986). Figure 10 compares the balloon and the SAGE II extinction profiles respectively at $.375 \mu \mathrm{~m}$ and $.385 \mu \mathrm{~m}$; figure 11 gives the same comparison for . $440 \mu \mathrm{~m}$ and $.453 \mu \mathrm{~m}$; the error due to the small difference in wavelength is in both cases smaller than $5 \%$ and we have not found necessary to introduce a correction. Figure 12 presents the comparison of the balloon and the SAGE II extinction profiles at $.84 \mu \mathrm{~m}$. The SAGE II profile at $1.02 \mu \mathrm{~m}$ has been converted to $.84 \mu \mathrm{~m}$ using a LND model ( $\mathrm{veff}_{\mathrm{eff}}=.25$ ) with an effective radius of $.10 \mu \mathrm{~m}$ above 22 km , and . $35 \mu \mathrm{~m}$ below 22 km , i.e. the model which gives in the best agreement between lidar and SAGE II profiles (section 2, figure 6); however the conversion between $1.02 \mu \mathrm{~m}$ and $.84 \mu \mathrm{~m}$ is not sensitive to the choice of the model as mentionned above. The balloon profiles exhibit oscillations which are smoothed on the SAGE II profiles. Above 21 km , the general agreement is very good up to $26-28 \mathrm{~km}$ for the $.44 / .453 \mu \mathrm{~m}$ and the $.375 / .385 \mu \mathrm{~m}$ profiles, and up to


Fig. 8. - Comparison of aerosol extinction profiles from SAGE II 11-10-84, 1627 GMT, $37^{\circ} 80 \mathrm{~N}, 5{ }^{\circ} 27 \mathrm{E}$ (dash-dotted line) and SAGE 11-11-84, 1639. CMT, $40^{\circ} 50 \mathrm{~N}, 0^{\circ} 56 \mathrm{E}$ (dashed-line) and from balloon limb photographs, 11-10-84, sunset, launch site $44^{\circ} \mathrm{N}$, $0^{\circ}$ (solid line); SAGE II $\lambda=.453 \mu \mathrm{~m}$; balloon $\lambda=.44 \mu \mathrm{~m}$.


Fig. 9.- Same as figure 8. $\lambda=1.02 \mu \mathrm{~m}$. Balloon data are converted from $.84 \mu \mathrm{~m}$ to $1.02 \mu \mathrm{~m}$ with LND aerosol models, $\mathrm{v}_{\text {eff }}=.10 ; \mathrm{r}_{\text {eff }}=$ $.28 \mu \mathrm{~m}$ (dashed-line), $r_{\text {eff }}=.05 \mu \mathrm{~m}$ (solid-line).


Fig.10.- Comparison of aerosol extinction profiles from SAGE II 04-22-85, $1906 \mathrm{GMT}, 47^{\circ} 17 \mathrm{~N}, 2^{\circ} 90 \mathrm{~W}$ (dashed line) and from balloon limb photographs, 04-22-85, sunset, launch site $44^{\circ} \mathrm{N}$, $0^{\circ}$ (solid line); SAGE II $\lambda=.385 \mu \mathrm{~m}$; balloon $\lambda=.375 \mu \mathrm{~m}$.




Fig.12.- Same as figure 10. $\lambda=.84 \mu \mathrm{~m}$. SAGE II data are converted from $1.02 \mu \mathrm{~m}$ to $.84 \mu \mathrm{~m}$ with LND aerosol models, $\mathrm{v}_{\text {eff }}=.25 ; \mathrm{r}_{\mathrm{eff}}=$ $.35 \mu \mathrm{~m}, \mathrm{z}<22 \mathrm{~km} ; \mathrm{r}_{\text {eff }}=.10 \mu \mathrm{~m}, \mathrm{z}>22 \mathrm{~km}$. Black circle is from balloon extinction.

32 km for the $.84 \mu \mathrm{~m}$ profile. Below 20 km , the balloon profiles deduced from scattering could be of poorer quality. However in this case the photographically measured extinction becomes significant and hence reliable. The value of aerosol extinction (total extinction minus Rayleigh and $O_{3}$ extinction) deduced from the balloon data agrees well with SAGE II results at 18 km altitude (Figure 3-12). Above 26 km the error bars on SAGE II profiles become very large for the short wavelength channels and the oscillations of the balloon profiles increase towards high altitudes and short wavelengths. However the balloon extinctions for . $44 \mu \mathrm{~m}$ and $.375 \mu \mathrm{~m}$ are systematically higher than the SAGE II extinctions above 26 km , which would point to smaller particles observed by the balloon.

The balloon limb photographs on April 22, 1985, provide an almost direct comparison with SAGE II profiles, for a close coincidence, and in a stable situation. The agreement for the three wavelengths is very good between 21 and 26 km .

## 4. BALLOON POLARIMETRIC OBSERVATIONS

The instrument (Herman et al, 1986) is a narrow field of view polarimeter operating at two wavelengths in the near infrared, $85 \mu \mathrm{~m}$ and $1.65 \mu \mathrm{~m}$. The scanning is performed in an horizontal plane by rotation of the gondola. Measurements of the radiance $L$ and of the degree of polarization $P$ at the two wavelengths can be made during the ascent and the descent, or at the ceiling level; the best conditions are sunset or sunrise, when the sun is close to the horizon, allowing the scattering angle $\theta$ to vary between $0^{\circ}$ and $180^{\circ}$. The data are first corrected for multiple scattering and for the reflection by the ground or by the clouds, in the case when the sun is above the horizon; the radiance is more conveniently expressed as a reflectance $\rho=\pi L / E$, where $E$ is the solar irradiance. The inversion procedure uses first the polarization $P\left(\theta_{0}\right)$ at $1.65 \mu \mathrm{~m}\left(\theta_{0}=100^{\circ}\right)$, assuming that the molecular contribution is negligible; this defines a family of LND size distribu-
tions ( $r_{m}$ versus $\sigma$ ); the reflectances $\rho\left(\theta_{0}\right)$ are used to retrieve the tangent optical depths for the two wavelengths; finally the polarization $P\left(\theta_{0}\right)$ at $.85 \mu \mathrm{~m}$ is used to select a model amongst the family found previousiy, It ls checked that the reflectance $\rho(\theta)$ and polarization $P(\theta)$ computed with this model agree with the measured values for the two wavelengths and all the scattering angles.

Four flights took place, one for ach period during the SAGE II European correlative programm. All the flights were launched from the CNES (Centre National d'Etudes Spatiales) Center at Aire sur l'Adour in the South West of France. The four flights are briefly described here and the results will be presented in section $5-2$.

On November 10,1984 (sunset flight), the aerosol layer was very inhomogeneous and unstable around the balloon. The data recorded between 14 km and 30 km , with a gap due to transmission problems between 19 km and 23 km , are therefore of poor quality. The polarization diagrams can be inverted only around $16-19 \mathrm{~km}$.

On November 28,1984 (sunrise filight), the conditions were better. Unfortunately the balloon did not fly above 24 km , but good quallity data were recorded between 14 km and 24 km .

During the filght of April 22, 1985, (sunset), the instrument broke down at celling level, and data were recorded oniy during the ascent between $15-22 \mathrm{~km}$ and $27-30 \mathrm{~km}$ when the gondola was not very stable and the sun was still rather high above the horlzon.

The flight of Ootober 12. 1985, took place during sunset in good stable conditions, and data were recorded from 16 to 33 km . However above 2 km , the aerosol content was low and the results are of better quality at low altitudes.
5. ANALYSIS OF THE RESULTS - VALIDATION OF THE THREE SHORT WAVELENGTH

CHANNELS

The aerosol extinction coefficient is retrieved from SAGE II data at four wavelengths $1.02 \mu \mathrm{~m}, .525 \mu \mathrm{~m}, .453 \mu \mathrm{~m}, .385 \mu \mathrm{~m}$, leading to a spectral extinction curve $\sigma^{\text {aer }}(\lambda)$ which could "in principle" be inverted to give the size distribution $n(r)$. The lidar backscattering profiles have been converted into extinction profiles at $1.02 \mu \mathrm{~m}$ using the aerosol model with $v_{\text {eff }}=.25$ which gives the best agreement to SAGE II profiles. Of course varying $v_{\text {eff }}$, within a reasonable range, leads to a family of size distributions characterized by ( $v_{\text {eff }}, r_{\text {eff }}$ ), which give the same conversion factor from lidar into extinction profiles. The model used for lidar conversion must be consistent with the model fitting the SAGE II spectral extinction $\sigma^{\text {aer }}(\lambda)$. The balloon polarization measurements lead to a retrieval of the size distribution $n(r)$, which best fits the polarization and the reflectance diagrams at . $85 \mu \mathrm{~m}$ and $1.65 \mu \mathrm{~m}$. This has also to be consistent with the SAGE II spectral extinction and with the lidar conversion factor. Finally the balloon limb photographs provide profiles to be compared to the SAGE II short wavelength extinction profiles (see Figures 8, 10, 11).

Inverting the SAGE II spectral extinction $\sigma^{\text {aer }}(\lambda)$ is a rather delicate problem and various approaches have been tried in order to retrieve two parameters of the size distribution, i.e. the effective radius $r_{\text {eff }}$ and the effective variance $v_{\text {eff }}$, or the mean radius $r_{m}$ and the variance $\sigma$. The discussion of this inversion problem is beyond the scope of the present work and will be left for a future contribution. We will limit ourselves here to deducing the effective radius $r_{\text {eff }}$ (Lenoble and Brogniez, 1985) for an arbitrary fixed variance ( $\mathrm{veff}_{\text {f }}=.25$ ), from the ratio $\sigma^{\text {aer }}(.453) / \sigma^{a e r}(1.02)$, or more conveniently from the related mean Angström coefficient $\alpha$ for the spectral interval . 453/1.02 $\mu \mathrm{m}$, defined by

$$
\begin{equation*}
\sigma^{\text {aer }}(\lambda)=\sigma^{\text {aer }}(1.02) \lambda^{-\alpha} \tag{7}
\end{equation*}
$$

Preliminary tests with a two parameter retrieval procedure suggests that the effective variance is generally smaller than .25 at the low levels, below 22 km (Brogniez and Lenoble, 1987).
5.1. Consistency of lidar conversion factor with SAGE II spectral extinction

For the period November $10-13,1984$, a stable situation is found only on November 13, when a lidar comparison was possible (Fig. 2). If we look at the Angström coefficient a for the wavelength interval $(.453 / 1.02 \mu \mathrm{~m})$, from the SAGE II profiles on November 13 , it varies approximately from .6 to 1.8 when the altitude increases from 15 km to 25 km , pointing to a decrease of the particle effective radius from about $.40 \mu \mathrm{~m}$ to $.20 \mu \mathrm{~m}$ with altitude; the variation of $\sigma^{\text {aer }}(\lambda)$ between $.525 \mu \mathrm{~m}$ and $.385 \mu \mathrm{~m}$ suggests, at least for the low altitudes, a rather small effective variance, around .1 or a little larger than . 1 . As mentioned above, the conversion factor from lidar backscattering into extinction is not very sensitive to the model for particles with reff larger than $.20 \mu \mathrm{~m}$ and the agreement found between the lidar and the SAGE II profiles on Figure 2 , would remain had we used the aerosol models derived from the SAGE II spectral extinction instead of the model $r_{\text {eff }}=.35 \mu \mathrm{~m}, \mathrm{v}_{\text {eff }}=.25$.

For the period November 26-30, 1984, the comparison between SAGE II and lidar profiles on November 27 (Fig. 3) requires an aerosol model with $r_{\text {eff }}=.25 \mu \mathrm{~m}$ for $z<21 \mathrm{~km}<z<26 \mathrm{~km}$, and $r_{\text {eff }}=.05 \mu \mathrm{~m}$ for $z>26 \mathrm{~km}$. For the low altitudes, the SAGE II spectral extinction suggests $r_{\text {eff }}$ between $.35 \mu \mathrm{~m}$ and $.24 \mu \mathrm{~m}$ ( $\alpha$ between . 8 and 1.5), which is consistent with the model ( $r_{\text {eff }}=.25 \mu \mathrm{~m}$ ) choosen for the lidar conversion, considering the small sensitivity of the conversion factor in this size range. However at higher altitudes, $\alpha$ increases from about 1.3 to 1.9 , which means a decrease of reff from about $.25 \mu \mathrm{~m}$ to $.18 \mu \mathrm{~m}$. Small particles as chooser for the lidar conversion factor would give $\alpha$ around 3 ; this is absolutely inconsistent with the SAGE II extinctin values in the short wavelength channels, which are much too low.

On November 28, as mentioned previously, the two SAGE II profiles are different above 25 km , but quite close between 16 km and 25 km , with a small extinction peak around 21 km for the western profile, which does not appear on the eastern profile (Fig. 4); in the peak the particles are slightly larger ( $\alpha=1.2$ for the western profile, instead of 1.4 at the same level in the eastern profile). The two lidar profiles (Frascatti and OHP) are rather different and the OHP profile agrees better with the SAGE II profiles. From SAGE II extinction profiles, $\alpha$ increases from .65 to 1.8 between 14 km and 24 km , which means $r_{\text {eff }}$ decreasing from . $40 \mu \mathrm{~m}$ to $.20 \mu \mathrm{~m}$; this is again consistent with the choice $r_{\text {eff }}=.35 \mu \mathrm{~m}$ for the lidar conversion factor.

On November 30 (Fig. 5) we have the closest coincidence between a lidar and a SAGE II observation. The conversion factor for an aerosol model with $r_{\text {eff }}=.35 \mu \mathrm{~m}$ gives a very good agreement between 13 km and 22 km (upper limit of the lidar profile). In this altitude range, the SAGE II Angström coefficient $\alpha$ varies from . 6 to 1.4 , which corresponds to particles with $r_{\text {eff }}$ decreasing slightly from about. $40 \mu \mathrm{~m}$ to $.25 \mu \mathrm{~m}$. This is again perfectly consistent with the choice of $\mathrm{r}_{\text {eff }}=.35 \mu \mathrm{~m}$ for the conversion of lidar data.

In April 1985, the aerosol layer over Europe during the observation period was very homogeneous and stable. The comparison between the SAGE II average profile and the Garmisch-Partenkirchen profile on April 21 was made using for the conversion factor $r_{\text {eff }}=.35 \mu \mathrm{~m}$ below 22 km and $\mathrm{r}_{\text {eff }}=.10 \mu \mathrm{~m}$ above. The SAGE II spectral extinction gives $r_{\text {eff }}$ decreasing from about. $35 \mu \mathrm{~m}$ at 13 km , to $.25 \mu \mathrm{~m}$ at 22 km , then to $.17 \mu \mathrm{~m}$ at 30 km (respectively $\alpha=.9, \alpha=1.4$ and $\alpha=2.0$ ). This is consistent with the choice $r_{\text {eff }}=.35 \mu \mathrm{~m}$ for the lidar conversion factor below 22 km , but not at higher altitudes, where the SAGE II spectral extinction leads to participate much larger than the particles which are found necessary in order to obtain a good agreement between the lidar an the SAGE II profile. We find again the same difficulty, as on November 27 : the aerosol model derived from SAGE II spectral extinction would lead to a poor agreement with the lidar profile at high altitudes;
whereas the agreement is obtained assuming much smaller particles than given by SAGE II short wavelength ohannels. The balloon limb photographs (section 3) on April 22 give extinction coefficlents larger than SAGE II above 25 km for the short wavelengths (Eig. 10 and 11), pointing to particles smalien than those retrieved from SAGE İ. But it is difficult with the rapid oscillations of the balloon profiles to deduce the spectral variation $\sigma^{a e r}(\lambda)$ at a glven leval and to make a quantitative comparison with SAGE 11 size distribution:

On october 12, 1985, the comparison between the SAGE il profile and the OHP lidar was made with $\mathrm{reff}_{\mathrm{ef}} \leq .17 \mu \mathrm{~m}$, between 16 km and 25 km , whereas the SAGE IX spectral extinction leads to reff decreasing from .34 . m to . $21 \mu \mathrm{~m}$ with altitude (a between .9 and 1.7 ). It is the only case, where we find some inconslstency between the best cholce for the lidar conversion factor and the best fit to aer $(\lambda)$ at low altitudes. On Figure 7, we have also drawn the extinction profile deduced from the lidar profile, using the aerosol model; which fits both the polarization measurements (see discussion in section $5-\bar{e}$ ) and the SAGE II spectral. extinction. The agreement with the SAGE il profile is definitely not as good as obtained with the model $v_{\text {eff }} \equiv .25, ~ r e f f=.17 \mu m$, but the disagreement appears only below 20 km and remains rather small : it might be attributed to smali local or temporal variation of the aerosol, as the observations are not exactiy colneident either in location, or in time.

These results are sumarized in table 1.

### 5.2. Balloon polarimetric observations and size distribution

The báloon polarimetric observations provide radiance and polarization diagrams at $.85 \mu \mathrm{~m}$ and $1.65 \mu \mathrm{~m}$, and their inversion leads to the petrieval of two parameters of the size distribution, assumed to be log-normal. However it must be kept in mind that the actual aerosol size distribution may not be close to lognormal and may not even be
monomodal. The retrieved size distribution must be understood as one of the many size distributions which give a good fit to the radiance and to the polarization of the diffuse radiation in the near infrared. The inversion of the SAGE II spectral extinction between $.385 \mu \mathrm{~m}$ and $1.02 \mu \mathrm{~m}$ is subject to the same remark than the inversion of the polarimetric data; the retrieved size distribution is one of many which give a good fit to the extinction coefficient in the visible range. Therefore, using the balloon polarimetric data to validate the SAGE II short wavelength channels is a rather delicate task, and the results must be considered with caution.

For the flight of November 10, 1984, figure 13 compares the tangent optical depth at $1.02 \mu \mathrm{~m}$ observed trom SAGE II on November 10 and 11 with the tangent optical depth observed by the balloon instrument. As noted previously, the SAGE II event tangent point is closer to the balloon launch site on 11-10 than on 11-11. The balloon data exhibit strong oscillations and have been averaged over 1 km . The balloon data exhibit strong oscillations and have been averaged over 1 km . The balloon optical depth values at $.85 \mu \mathrm{~m}$ have been converted into values at $1.02 \mu \mathrm{~m}$, using an aerosol effective radius of $.28 \mu \mathrm{~m}$ below 20 km and of $.10 \mu \mathrm{~m}$ above 22 km ; the influence of the model choice is however small. The balloon tangent optical depth profile given by the polarimeter is closer to the SAGE II profile on November 10, whereas the extinction profile deduced from the balloon limb photographs on the same day was closer to the SAGE II profile on November 11. This is not too surprising in a very unstable situation, as the two balloons were not operating exactly at the same place and at the same time. The complete inversion of the polarization data has been performed only for the altitude range $16-19.5 \mathrm{~km}$; the retrieved size distribution has an effective radius $r_{\text {eff }}=.35 \mu \mathrm{~m}$ and an effective variance $v_{\text {eff }}=.17$ between 16 and 17 km ; between 17.5 and 19.5 km , the effective radius is slightly smaller $r_{\text {eff }}=.29 \mu \mathrm{~m}$ with $v_{\text {eff }}=.14$. This is in excellent agrement with the size distribution retrieved from the SAGE II extinction ratio $\sigma^{\text {aer }}(.45) / \sigma^{\text {aer }}(1.02)$, which gives for November 11 , reff


Fig.13. - Comparison of the tangent optical depth at $1.02 \mu \mathrm{~m}$ measured by SAGE II on 11-10-84, $1627 \mathrm{GMT}, 37^{\circ} 80 \mathrm{~N}, 5^{\circ} 27 \mathrm{E}$, (solid line) and on 11-11-84, $16390 \mathrm{GMT}, 40^{\circ} 50 \mathrm{~N}, 56^{\circ} \mathrm{E}$ (dashed line), and measured by the balloon polarimeter on 11-10-84, sunset, launch site $44^{\circ} \mathrm{N}, 0^{\circ}$ (crosses). The conversion of the balloon data from $.85 \mu \mathrm{~m}$ to $1.02 \mu \mathrm{~m}$ has been done with $\mathrm{r}_{\text {eff }}=.28 \mu \mathrm{~m}$ below 20 km and $\mathrm{r}_{\text {eff }}=.10 \mu \mathrm{~m}$ above 22 km .
decreasing from $.32 \mu \mathrm{~m}$ to $.23 \mu \mathrm{~m}$ if we assume $\mathrm{v}_{\text {eff }}=.25$ and from $.36 \mu \mathrm{~m}$ to $.27 \mu \mathrm{~m}$ for $\mathrm{v}_{\mathrm{eff}}=.10$, between 15 km to 20 km . For November 10 , SAGE II data give, for the same altitude range $15-20 \mathrm{~km}$, an almost constant effective radius $r_{\text {eff }}=.30 \mu \mathrm{~m}$, if we assume $\mathrm{v}_{\text {eff }}=.25$, and $r_{\text {eff }}=$ $.33 \mu \mathrm{~m}$ with $\mathrm{v}_{\mathrm{eff}}=.10$.

The flight of November 28, 1984, took place in more stable conditions. The tangent optical depth measured by the balloon above 20 km is larger by about a factor two, than the optical depth measured by SAGE II, whereas at the lower levels ( $13-17 \mathrm{~km}$ ) the two values agree reasonably well. No explanation has been found for this disagreement, which may just be due to local conditions. The polarization data lead to a size distribution with an effective radius almost constant around $.22 \mu \mathrm{~m}$, and an effective variance decreasing from .80 to . 18 , between 15 km and 22 km . The SAGE II extinction ratio $\sigma^{\operatorname{aer}}(.45) / \sigma^{\text {aer }}(1.02)$ leads to an effective radius decreasing from $.38 \mu \mathrm{~m}$ to $.22 \mu \mathrm{~m}$, assuming $\mathrm{v}_{\text {eff }}=$ .25, for the same altitude range. The large variance found by the balloon at low levels seems to confirm the presence of particles different from those observed by SAGE II.

Unfortunately on April 22, 1985, no inversion of the polarization diagram was possible, due to the instability of the data. However at a few levels, a relative stabilization appeared, and the diagram can be used for direct comparisons. Figure 14 shows the polarization diagram for the two wavelengths (.85 $\mu \mathrm{m}$ and $1.65 \mu \mathrm{~m})$ and three altitude levels $(15 \mathrm{~km}, 18.2 \mathrm{~km}$ and 21.5 km$)$; the dots are the experimental data and present a rather large dispersion. The solid lines show the polarization computed with models derived from a best fit to the SAGE II spectral extinction ( $v_{\text {eff }}=.17 ; r_{\text {eff }}=.38 \mu \mathrm{~m}$ at 15 km and $r_{\text {eff }}=.29 \mu \mathrm{~m}$ at 18.2 km and 21.5 km ). The comparison is satisfying. Unfortunately no such comparison was possible at higher levels.

The flight of October 12, 1985 provides another good comparison to SAGE II data. Figure 15 presents the vertical profiles of $r_{\text {eff }}$ and $v_{\text {eff }}$ retrieved from the polarization data. Above 22 km , the effective



Fig.15. - Vertical profiles of the effective radius (upper curve) and of the effective variance (lower curve) of the aerosol size distribution retrieved from the balloon polarimetric data on October 12, 1985.
variance increases rapidly and stabilizes around .9, whereas the effective radius presents very large oscillations; these results at high altitudes are certainly dubious, because the aerosol content becomes very low above 22 km and the signal to noise ratio becomes bad. However the large value retrieved for $v_{\text {eff }}$ could suggest that the size distribution becomes bimodal at high altitudes; therefore the retrieval procedure, which assumes a monomodal distribution, leads to erratic results. Figure 16 demonstrates for 17.5 km the good quality of the inversion. Figure 17 shows the tangent optical depth at $1.02 \mu \mathrm{~m}$ deduced from the measured optical depth at $.85 \mu \mathrm{~m}$, using LND models which incorporate, at each altitude, the effective radius and the effective variance retrieved from the polarization data and averaged over 1 km ; it is compared with the SAGE II tangent optical depth at $1.02 \mu \mathrm{~m}$ on October 12, 1985, morning, at $7^{\circ} \mathrm{W}$ of the launch site. The similarity, above 16 km , between the four SAGE II profiles over the zone on October 12 and 13, justify the comparison, despite the not very close coincidence in time and location. Figure 18 compares the extinction ratios for the three short wavelengths $\sigma^{\operatorname{aer}}(.525) / \sigma^{\operatorname{aer}}(1.02)$, $\sigma^{\text {aer }}(.45) / \sigma^{\text {aer }}(1.02), \sigma^{\text {aer }}(.385) / \sigma^{\text {aer }}(1.02)$, measured by SAGE II, and computed at each level with the size distribution retrieved from the polarization data and averaged over 1 km . As a result of the low aerosol content the extinction ratios derived from the polarization data are somewhat inaccurate. However the agreement is good.

## 6. CONCLUSIONS

Although the SAGE II $1.02 \mu \mathrm{~m}$ channel allows retrieval of the extinction profile at very low levels, only the profiles a few km above the tropopause ( $12 \mathrm{~km}-15 \mathrm{~km}$ ) and higher have been considered; at lower altitudes, the variability is such that only almost coincident observations would be necessary to validate SAGE II profiles.

From the data and the discussions presented in the previous sections, we must consider separately two altitude ranges. For safety, we will refer them as below 23 km and above 25 km , being understood that


Fig.16. - Test of the inversion of the balloon polarimetric data on October 12, 1985. The dots are the measured values which are to be compared with the curves computed with the retrieved aerosol model. Reflectance at $.85 \mu \mathrm{~m}$ (upper left), and $1.65 \mu \mathrm{~m}$ (upper right) and degree of polarization in percent at $.85 \mu m$ (lower left) and $1.65 \mu \mathrm{~m}$ (lower right) versus the scattering angle.


Fig.17.- Comparison of the tangent optical depth at $1.02 \mu \mathrm{~m}$ measured by SAGE II on 10-12-85, $0639 \mathrm{GMT}, 41^{\circ} 25 \mathrm{~N},-7 \circ 10 \mathrm{E}$ (solid line) and measured by the balloon polarimeter on 10-12-85 sunset, launch site $44^{\circ} \mathrm{N}, 0^{\circ}$ (black circles). The conversion of the balloon data at $.85 \mu \mathrm{~m}$ to $1.02 \mu \mathrm{~m}$ has been made with the aerosol model retrieved from the polarization data.



F1g.18.- Comparison of the aerosol extinction ratio $\sigma^{\text {aer }}(\lambda) / \sigma^{\text {aer }}(1.02)$ measured by SAGE if on 10-12-85, $0639 \mathrm{GMT}, 41^{\circ} 25 \mathrm{~N},-7010 \mathrm{E}$ (solid line) and computed uslng the aerosol model retrieved from the balloon polarization data on 10-12-85, sunset, launch site $44^{\circ} \mathrm{N}, 0^{\circ}$ (black circles); upper curves, left $\lambda=.525 \mu \mathrm{~m}$, right $\lambda=.453 \mu \mathrm{~m}$; lower curve $\lambda=.385 \mu \mathrm{~m}$.
the cutoff between the two ranges is somewhat variable, depending on the events and on the kind of observations.

Below 23 km , we have at our disposal a large series of data, including lidar profiles (both ruby and $N d Y a g$ ), limb photographs and polarization diagrams. The main conclusions are as follows :

- the SAGE II extinction profiles at $1.02 \mu \mathrm{~m}$ agree within the error bars with the several extinction profiles deduced from the lidar backscattering profiles using a conversion factor, consistent with the SAGE II spectral variation of the extinction coefficient (Figures 2 to 7 and Table 1). These comparisons comprise one case (November 30, 1984) of very close coincidence in time ( 4 hours) and in location ( 100 km ) between the lidar and the SAGE II observations, and several cases with a very stable and homogeneous aerosol. layer, as proved by the comparisons between various SAGE II and lidar profiles over Europe for the experiment. period;
- however the consistency of the choosen backscatter into extinction conversion factor with the aerosol size distribution, retrieved from the four wavelength SAGE II extinction, does not really validate the SAGE II four channels because the conversion factor is almost insensitive to the aerosol model, as long as the effective radius is larger than $.20 \mu \mathrm{~m}$, which is the case in this altitude range;
- the SAGE II extinction profile at $1.02 \mu \mathrm{~m}$ also agrees with the limb photographs profile at $.84 \mu \mathrm{~m}$, the conversion between $1.02 \mu \mathrm{~m}$ and $.84 \mu \mathrm{~m}$ being only very slightly sensitive to the aerosol model; the agreement is particularly good on April 22, 1985 (Figure 12) when the conditions are quite stable and the coincidence very close;
- the SAGE II tangent optical depth profiles at 1.02 , 1 m generally agrees within the error bars with the optical depth profiles obtained by the balloon borne polarimetric instrument;

TABLE 1 Comparison of the aerosol effective radius $r_{\text {eff }}$ used for lidar/SAGE II best fit at $1.02 \mu \mathrm{~m}$ and retrieved from SAGE II extinction ratio $\sigma^{\text {aer }}(.45) / \sigma^{\text {aer }}(1.02)$.

| Date | Altitude (km) | $\begin{gathered} r_{\text {eff }}\left({ }^{(l i d a r)}\right. \end{gathered}$ | $\begin{gathered} \mathrm{r}_{\mathrm{eff}}(\text { SAGE II) } \\ (\mu \mathrm{m}) \end{gathered}$ |
| :---: | :---: | :---: | :---: |
| 11-13-84 | 12-25 | . 35 | . 40 to .20 |
| 11-27-84 | 12-21 | . 25 | . 35 to . 24 |
|  | 21-26 | . 10 | . 24 to . 18 |
|  | 26-30 | . 05 | . 18 |
| 11-28-84 | 12-23 | . 35 | . 40 to . 20 |
| 11-30-84 | 13-32 | . 35 | . 40 to .25 |
| 04-21-85 | 10-22 | . 35 | .35 to .25 |
|  | 22-30 | . 10 | . 25 to .17 |
| 10-12-85 | 16-25 | . 17 | . 34 to . 21 |

- the SAGE II extinction profiles at. $453 \mu \mathrm{~m}$ and $.385 \mu \mathrm{~m}$ have been compared directly to the extinction profiles at about the same wavelengths deduced from the limb photographs. April 22, 1985 (Figures 10 and 11) corresponds to a close coincidence (sunset, less than 100 km between the two observations). Despite the oscillations revealed by the high resolution of the photographs, the general agreement between the balloon and the SAGE II profiles is a good validation of the two channels $.453 \mu \mathrm{~m}$ and $.385 \mu \mathrm{~m}$, for the altitude range $21-25 \mathrm{~km}$;
- the size distributions derived from the balloon polarization measurements and from SAGE II spectral extinction generally agree; the extinction ratio profils $\sigma^{\text {aer }}(\lambda) / \sigma^{\text {aer }}(1.02)$ at $.525 \mu m, 453 \mu \mathrm{~m}$ and $.385 \mu \mathrm{~m}$ computed with the size distribution retrieved from the polarlzation measurements un Ocluber 12, 1985 agree well within the error bars with the corresponding SAGE II profiles (Figure 18). This is again a satisfying validation of the SAGE II short wavelength channels.

Above 25 km , the situation is not as good. Most instruments failed in observing the low content of aerosols at these altitudes and only a few data remain available : ruby lidar profiles on November 27, 1984 and April 21, 1985, limb photograph profiles on April 22, 1985; even these available data are not of the same quality as at lower levels. On the other hand, whereas the SAGE II profile at $1.02 \mu \mathrm{~m}$ remains rahter good up to 30 km , the three short wavelength profiles have increasingly large error bars above 25 km . The main conclusions for the high altitude range are the following ;

- the extinction profiles at $1.02 \mu \mathrm{~m}$ deduced from the lidar backscattering profiles can be put into agreement with the SAGE II profiles (Figures 3 and 6) using for the conversion of backscattering into extinction an aerosol model with very small particles ( $r_{\text {eff }}=.10 \mu \mathrm{~m}$ to $.05 \mu \mathrm{~m}$ );
- this choice is inconsistent with the size distribution derived from the SAGE II spectral variation of extinction, which leads to $r_{\text {eff }} \simeq$ $.18 \mu \mathrm{~m}$. In this size range the conversion factor of backscattering into
extinction is very sensitive to the aerosol model, and choosing $r_{\text {eff }}=$ $.18 \mu \mathrm{~m}$ would destroy the agreement of the SAGE II and the lidar profiles on Figures 3 and 6;
- the limb photography profile at . $84 \mu \mathrm{~m}$ agrees perfectly well (Figure 12) with the $1.02 \mu \mathrm{~m}$ SAGE II profile converted at $.84 \mu \mathrm{~m}$ (conversion not very sensitive to the model choice);
- the limb photography profiles at the short wavelengths show extinction significantly larger than SAGE II (Figures 10 and 11). This suggests particles with $r_{\text {eff }}$ smaller than $.18 \mu \mathrm{~m}$, but the very large oscillations of the profiles do not allow a retrieval of $r_{e f f}$.

Whereas a good validation of SAGE II aerosol extinction profiles is obtained below 23 km , it seems difficult to draw a clear conclusion from the few observations above 25 km . It is likely that the SAGE II $1.02 \mu \mathrm{~m}$ profile, which has small error bars, remains good. But the three SAGE II short wavelength channels, as well as the ruby lidar profile and the limb photographs, have very large uncertainties at these high levels; it is hard to decide what must be better believed.

A very tentative guess to explain at least a part the contradiction at high altitudes is that the size distribution becomes bimodal; for a fixed value of the ratio $\sigma^{\operatorname{aer}}(.45) / \sigma^{\text {aer }}(1.02)$, it has been shown (Lenoble and Brogniez, 1985) that the lidar conversion factor generally increases when a second mode is added to a size distribution. Qualitatively this could reconciliate the SAGE II spectral extinction. with the choice of the conversion factor necessary to have agreement between the lidar and the SAGE II $1.02 \mu \mathrm{~m}$ profile. This could also explain the bad quality of the polarization data inversion and the rapid increase of $v_{\text {eff }}$ above 23 km .

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